Key for Periodic Trends in Properties of Elements I

The following figure shows the first ionization energies, IE_1 , for the elements $H \longrightarrow Ca$ as a function of their atomic numbers, Z.



atomic number

We know that IE_1 is proportional to Z, and yet the data do not show a steady increase in IE_1 as a function of Z. For each decrease in IE_1 that you see in this figure—for example, there is a decrease in IE_1 from helium to lithium—explain the source of the decrease.

- The decrease from He to Li, from Ne to Na, or from Ar to K is explained by a change in the shell that holds the electron: from n = 1 to n = 2 for He and Li, from n = 2 to n = 3 for Ne and Na, or from n = 3 to n = 4 for Ar and K. An increase in n implies an increase in the distance of the electron from the nucleus, r. As IE is inversely proportional to r, this decreases IE_1 .
- The decrease from Be to B, or from Mg to Al is explained by a change in the subshell that holds the electron: from 2s to 2p for Be and B, or from 3s to 3p for Mg and Al. As an ns electron screens an np electron more than an np electron screens an ns electron, Z_{eff} is greater from an ns electron. As IE is directly proportional to Z_{eff} , this decreases IE_1 .
- The decrease from N to O, or from P to S is explained by how electrons enter orbitals of the same subshell. The first three electrons in the 2p orbital for N and the 3p orbital for P occupy separate orbitals; the additional 2p electron for O and 3p electron for S must pair up with another electron. The repulsion between the electrons makes them easier to remove, which decreases IE_1 .

The photoelectron spectrum for scandium, Sc, has seven peaks with the following ionization energies, IE, and relative heights:

| | Peak 1 | Peak 2 | Peak 3 | Peak 4 | Peak 5 | Peak 6 | Peak 7 |
|--|--|-------------------|---------------------|-------------------|---------------------|---------------------|--|
| $\overline{IE \text{ (MJ/mol)}}$ relative height orbital | $\begin{array}{c} 433\\2\\1s^2\end{array}$ | 48.5 2 $2s^2$ | $39.2 \\ 6 \\ 2p^6$ | 5.44 2 $3s^2$ | $3.24 \\ 6 \\ 3p^6$ | $0.77 \\ 1 \\ 3d^1$ | $\begin{array}{c} 0.63 \\ 2 \\ 4s^2 \end{array}$ |

Complete the table by completing the row giving the orbital designation for each peak. Explain how you made your decisions.

• The orbital designations are shown above. The only question to address is whether the last two peaks are $3d^1, 4s^2$ or $4s^1, 3d^2$. From Madelung's rule, we expect the electrons to fill the 4s orbital before they enter the 3d orbital; thus, we expect Sc to have two 4s electrons. Once filled, the electrons in the 4s orbital become somewhat easier to remove than a 3d electron. This is consistent with Slater's rules, which suggest that 3d electrons screen 4s electrons, but that 4s electrons do not screen 3d electrons.

The 1s, 2s, and 3s electrons in argon, Ar, have ionization energies of 309, 31.5, and 2.82 MJ/mol, respectively. What is the full electron configuration for Ar?

• The electron configuration for argon is $1s^22s^22p^63s^23p^6$.

We can use Slater's rules to estimate Z_{eff} for any electron in Ar. For example, for a 1s electron in Ar, we have

$$Z_{\text{eff}} = Z - S = 18 - (1)(0.35) = 17.65$$

Note that the electron of interest is a 1s electron, that there is one additional electron in the 1s orbital, which accounts for the term -(1)(0.35), and that we ignore the electrons in all other orbitals because they have smaller ionization energies. Using Slater's rules, estimate Z_{eff} for the 2s and the 3s electrons in Ar.

• Using Slater's rules, we first group the orbitals

$$(1s^2)$$
 $(2s^22p^6)$ $(3s^23p^6)$

For a 2s electron we have $Z_{\text{eff}} = Z - S = 18 - (5)(0.35) - (2)(0.85) = 14.55$, and for a 3s electron we have $Z_{\text{eff}} = Z - S = 18 - (5)(0.35) - (8)(0.85) - (2)(1.00) = 7.45$.

Based on your values for Z_{eff} , are the effective nuclear charges sufficient to explain the differences in the ionization energies for the 1s, 2s, and 3s electrons in argon? If no, then offer an additional explanation.

• The ionization energies show that it is approximately $10 \times$ harder to remove a 2s electron than a 3s electron (31.5 vs. 2.82 MJ/mol). The value of $Z_{\rm eff}$ for a 2s electron relative to a 3s electron, however, is less than $2 \times$ greater (14.55 vs. 7.45), which means we cannot explain the difference in their ionization energies based on $Z_{\rm eff}$ alone. It must be the case that the radius for a 2s electron is smaller than expected, suggesting that the greater the effective nuclear charge, the more the electrons are pulled toward the nucleus and the smaller the radius. A similar argument applies to the relative ionization energies of the 1s and 2s electrons.

An element's first ionization energy, IE_1 , is the energy needed to remove the least tightly-held electron, which we represent by the reaction $E \longrightarrow E^+ + e^-$. The second ionization energy, IE_2 is the energy needed to remove the next least tightly-held electron, which we represent by the reaction $E^+ \longrightarrow E^{2+} + e^-$. The third and subsequent ionization energies follow the same pattern. The first and second ionization energies for helium, He, are 2370 and 5250 kJ/mol, respectively. Given that both electrons are in the same orbital, why is IE_2 more than 2× than that for IE_1 ?

• When we remove the first electron from He, we leave behind an ion with a positive charge, He⁺. Three things make the second electron harder to remove. First, the two electrons in He screen each other from seeing the full charge of the nucleus, which means $Z_{\text{eff}} < 2$; once we remove one electron, the one that remains has $Z_{\text{eff}} = 2$. Second, the increase in Z_{eff} means the radius shrinks as the one remaining electron is pulled inward. And, third, it requires more energy to remove a negative charge from an ion that carries a positive charge than to remove it from a neutral atom.

The element beryllium, Be, has four electrons and four ionization energies, which are 900, 1760, 14850, and 21000 kJ/mol, respectively. As we expect, each ionization energy is greater than the preceding ionization energy. There is, however, a significant increase from IE_2 to IE_3 . Offer an explanation for this observation.

• The electron configuration for Be is $1s^22s^2$. The first and second electrons removed are from the 2s orbital. The third and fourth electron are in a 1s orbital and have much greater ionization energies because they are closer to the nucleus.